Hamiltonian geophysical fluid dynamics

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- Hamiltonian dynamics is a very beautiful, and very powerful, mathematical formulation of physical systems
 - All the important models in GFD are Hamiltonian
- Since it is a general formulation, it provides a framework for "meta-theories", providing traceability between different approximate models of a physical system
 - e.g. barotropic to quasi-geostrophic to shallow-water to hydrostatic primitive equations to compressible equations
 - Symmetries and conservation laws are linked by Noether's theorem
- In their pure formulation, Hamiltonian systems are conservative; but the Hamiltonian formulation provides a framework to understand forced-dissipative systems too
 - The nonlinear interactions are generally conservative
 - Example: energy budget (APE and Lorenz energy cycle)
 - Example: momentum transfer by waves

Hamilton's equations for a canonical system:

$$\frac{\mathrm{d}q_i}{\mathrm{d}t} = \frac{\partial \mathcal{H}}{\partial p_i}, \qquad \frac{\mathrm{d}p_i}{\mathrm{d}t} = -\frac{\partial \mathcal{H}}{\partial q_i} \qquad (i = 1, \dots, N)$$

 For a Newtonian potential system, we get Newton's second law:

$$\mathcal{H} = (|\mathbf{p}|^2/2m) + U(\mathbf{q}) \quad \Rightarrow \quad m \frac{\mathrm{d}^2 q_i}{\mathrm{d}t^2} = -\frac{\partial U}{\partial q_i} \qquad (i = 1, \dots, N)$$

 $= \frac{\partial \mathcal{H}}{\partial q_i} \frac{\partial \mathcal{H}}{\partial p_i} - \frac{\partial \mathcal{H}}{\partial p_i} \frac{\partial \mathcal{H}}{\partial q_i} = 0$

Conservation of energy follows: $\frac{d\mathcal{H}}{dt} = \frac{\partial \mathcal{H}}{\partial q_i} \frac{dq_i}{dt} + \frac{\partial \mathcal{H}}{\partial p_i} \frac{dp_i}{dt}$ (repeated indices summed)

Symplectic formulation:

$$\frac{\mathrm{d}u_i}{\mathrm{d}t} = J_{ij} \frac{\partial \mathcal{H}}{\partial u_j} \qquad (i = 1, \dots, 2N)$$

$$\mathbf{u} = (q_1, \dots, q_N, p_1, \dots, p_N)$$

$$J = \begin{pmatrix} 0 & I \\ -I & 0 \end{pmatrix}$$

- The symplectic formulation of Hamiltonian dynamics can be generalized to other J, which have to satisfy certain mathematical properties
- Among these is skew-symmetry, which guarantees energy conservation:

$$\frac{\mathrm{d}\mathcal{H}}{\mathrm{d}t} = \frac{\partial \mathcal{H}}{\partial u_i} \frac{\mathrm{d}u_i}{\mathrm{d}t} = \frac{\partial \mathcal{H}}{\partial u_i} J_{ij} \frac{\partial \mathcal{H}}{\partial u_j} = 0$$

 The canonical J is invertible. If J is non-invertible, then Casimirs are defined to satisfy

$$J_{ij}\frac{\partial \mathcal{C}}{\partial u_j}=0 \qquad (i=1,\ldots,2N)$$

Casimirs are invariants of the dynamics since

$$\frac{\mathrm{d}\mathcal{C}}{\mathrm{d}t} = \frac{\partial \mathcal{C}}{\partial u_i} \frac{\mathrm{d}u_i}{\mathrm{d}t} = \frac{\partial \mathcal{C}}{\partial u_i} J_{ij} \frac{\partial \mathcal{H}}{\partial u_j} = -\frac{\partial \mathcal{H}}{\partial u_i} J_{ij} \frac{\partial \mathcal{C}}{\partial u_j} = 0$$

- Example of a non-canonical Hamiltonian representation:
 Euler's equations for a rigid body. The dependent variables
 are the components of angular momentum about principal
 axes, and the total angular momentum is a Casimir invariant.
- Cyclic coordinates: e.g. rotational symmetry implies conservation of angular momentum

$$\frac{\partial H}{\partial q_i} = 0 \Rightarrow \frac{dp_i}{dt} = 0 \qquad \text{for a given } i$$

 More generally, the link between symmetries and conservation laws is provided by *Noether's theorem*:

Given a function $\mathcal{F}(\mathbf{u})$, define $\delta_{\mathcal{F}}u_i = \varepsilon J_{ij}(\partial \mathcal{F}/\partial u_j)$

Then
$$\delta_{\mathcal{F}}\mathcal{H} = \frac{\partial \mathcal{H}}{\partial u_i} \delta_{\mathcal{F}} u_i = \varepsilon \frac{\partial \mathcal{H}}{\partial u_i} J_{ij} \frac{\partial \mathcal{F}}{\partial u_j}$$

But

$$\frac{\mathrm{d}\mathcal{F}}{\mathrm{d}t} = \frac{\partial \mathcal{F}}{\partial u_i} \frac{\mathrm{d}u_i}{\mathrm{d}t} = \frac{\partial \mathcal{F}}{\partial u_i} J_{ij} \frac{\partial \mathcal{H}}{\partial u_j}$$

and hence $\delta_{\mathcal{F}}\mathcal{H}=0$ if and only if $d\mathcal{F}/dt=0$

- Casimir invariants are associated with 'invisible' symmetries since $\delta_{\mathcal{C}}\mathbf{u}=0$
- Example: rigid body
- Barotropic dynamics is a Hamiltonian system

$$\begin{split} \frac{\partial \omega}{\partial t} &= -\mathbf{v} \cdot \boldsymbol{\nabla} \omega = -\partial(\psi, \omega) \\ \delta \mathcal{H} &= \delta \int \int \frac{1}{2} |\boldsymbol{\nabla} \psi|^2 \, \mathrm{d}x \, \mathrm{d}y \\ &= \int \int \boldsymbol{\nabla} \psi \cdot \delta \boldsymbol{\nabla} \psi \, \mathrm{d}x \, \mathrm{d}y \\ &= \int \int \{ \boldsymbol{\nabla} \cdot (\psi \delta \boldsymbol{\nabla} \psi) - \psi \delta \omega \} \, \mathrm{d}x \, \mathrm{d}y \end{split} \qquad \text{(assuming boundary terms vanish)}$$

- Functional derivatives are just the infinite-dimensional analogue of partial derivatives; they can reflect non-local properties
- Barotropic dynamics can be written in symplectic form as:

$$\frac{\partial \omega}{\partial t} = J \frac{\delta \mathcal{H}}{\delta \omega}$$
 where $J = -\partial(\omega, \cdot)$

The Casimir invariants are:

$$C = \int \int C(\omega) dx dy$$
 with $\frac{\delta C}{\delta \omega} = C'(\omega)$

and correspond to Lagrangian conservation of vorticity

Symmetry in x and conservation of x-momentum:

$$-\varepsilon \frac{\partial \omega}{\partial x} = \delta_{\mathcal{M}} \omega = \varepsilon J \frac{\delta \mathcal{M}}{\delta \omega} = -\varepsilon \partial \left(\omega, \frac{\delta \mathcal{M}}{\delta \omega} \right)$$

$$\delta \mathcal{M}/\delta \omega = y. \qquad \mathcal{M} = \int \int y \omega \, \mathrm{d}x \, \mathrm{d}y = \int \int y \left(\frac{\partial v}{\partial x} - \frac{\partial u}{\partial y} \right) \, \mathrm{d}x \, \mathrm{d}y$$

$$Kelvin's \ impulse \qquad \qquad = \int \int u \, \mathrm{d}x \, \mathrm{d}y \qquad \qquad \text{(ignoring boundary terms)}$$

Similarly for y-momentum and angular momentum:

$$\mathcal{M} = -\iint x\omega \, dx \, dy = \iint \nu \, dx \, dy$$
$$\mathcal{M} = -\iint \frac{1}{2} r^2 \omega \, dx \, dy = \iint \hat{\mathbf{z}} \cdot (\mathbf{r} \times \mathbf{v}) \, dx \, dy$$

 Quasi-geostrophic dynamics is analogous; e.g. for continuously stratified flow

$$\mathcal{H} = \iiint \frac{\rho_s}{2} \left\{ |\nabla \psi|^2 + \frac{1}{S} \psi_z^2 \right\} dx dy dz$$

$$q(x, y, z, t) = \psi_{xx} + \psi_{yy} + \frac{1}{\rho_s} \left(\frac{\rho_s}{S} \psi_z \right)_z + f + \beta y$$

$$\delta \mathcal{H} = \left[\iiint \frac{\rho_s}{S} \psi \delta \psi_z dx dy \right]_{z=0}^{z=1}$$

$$+ \iiint \left\{ \nabla \cdot (\rho_s \psi \delta \nabla \psi) - \rho_s \psi \delta q \right\} dx dy dz$$

- Now in addition to potential vorticity q(x,y,z,t), we need to consider potential temperature on horizontal boundaries ψ_z(x,y,t) [and possibly also circulation on sidewalls]
- Note that for the QG model these quantities also evolve advectively, like vorticity in barotropic dynamics:

$$\frac{\partial \omega}{\partial t} = -\mathbf{v} \cdot \nabla \omega = -\partial(\psi, \omega)$$

Analogously, the Casimir invariants and x-momentum are:

$$C = \iiint C(q_z) \, dx \, dy \, dz$$

$$+ \iiint C_0(\psi_z) \, dx \, dy \Big|_{z=0} + \iiint C_1(\psi_z) \, dx \, dy \Big|_{z=1}$$

$$\mathcal{M} = \iiint \rho_s yq \, dx \, dy \, dz$$

$$+ \iiint \frac{\rho_s}{S} y\psi_z \, dx \, dy \Big|_{z=0} - \iiint \frac{\rho_s}{S} y\psi_z \, dx \, dy \Big|_{z=1}$$

Rotating shallow-water dynamics:

$$\frac{\partial \mathbf{v}}{\partial t} + (f\hat{\mathbf{z}} + \nabla \times \mathbf{v}) \times \mathbf{v} + \nabla \left(\frac{1}{2}|\mathbf{v}|^2\right) = -g\nabla h,$$

$$\frac{\partial h}{\partial t} + \nabla \cdot (h\mathbf{v}) = 0 \qquad \mathcal{H} = \iint \frac{1}{2} \{h|\mathbf{v}|^2 + gh^2\} \, \mathrm{d}x \, \mathrm{d}y$$

$$\frac{\partial \mathcal{H}}{\partial \mathbf{v}} = h\mathbf{v}, \qquad \frac{\partial \mathcal{H}}{\partial h} = \frac{1}{2}|\mathbf{v}|^2 + gh$$

$$J = \begin{pmatrix} 0 & q & -\hat{c}_x \\ -q & 0 & -\hat{c}_y \\ -\hat{c}_x & -\hat{c}_y & 0 \end{pmatrix} \qquad q = (f + \hat{\mathbf{z}} \cdot \nabla \times \mathbf{v})/h$$

$$\mathcal{M} = \iint h(u - fy) dx dy$$
 $C = \iint hC(q) dx dy$

- Disturbance invariants: arguably the most powerful application of Hamiltonian geophysical fluid dynamics
- Ambiguities about the energy of a wave...
- Ambiguities about the momentum of a wave...
- If *u*=*U* is a steady solution of a Hamiltonian system, then

$$J \frac{\delta \mathcal{H}}{\delta \mathbf{u}} \bigg|_{\mathbf{u} = \mathbf{U}} = 0$$

- For a canonical system, J is invertible so $\delta \mathcal{H}/\delta \mathbf{u} = 0$ at $\mathbf{u} = \mathbf{U}$.
 - Hence the disturbance energy is quadratic
- But for a non-canonical system, this is not true and the disturbance energy is generally linear in the disturbance
 - Not sign-definite
 - Cannot define stability, normal modes, etc.

Pseudoenergy:

$$J\frac{\delta\mathcal{H}}{\delta\mathbf{u}}\bigg|_{\mathbf{u}=\mathbf{U}} = 0 \quad \text{implies} \quad \frac{\delta\mathcal{H}}{\delta\mathbf{u}}\bigg|_{\mathbf{u}=\mathbf{U}} = -\frac{\delta\mathcal{C}}{\delta\mathbf{u}}\bigg|_{\mathbf{u}=\mathbf{U}} \quad \text{for some Casimir } C$$
Thus $\delta(\mathcal{H}+\mathcal{C}) = 0$ at $\mathbf{u}=\mathbf{U}$.

$$\mathcal{A} = (\mathcal{H} + \mathcal{C})[\mathbf{u}] - (\mathcal{H} + \mathcal{C})[\mathbf{U}]$$
 is then both conserved and quadratic in the disturbance

Example: Available potential energy (APE) for the 3D stratified Boussinesq equations

$$\mathcal{H} = \int \int \left\{ \frac{1}{2} \rho_{\rm s} |\mathbf{v}|^2 + \rho gz \right\} dx dy dz \qquad \frac{\delta \mathcal{H}}{\delta \mathbf{v}} = \rho_{\rm s} \mathbf{v}, \qquad \frac{\delta \mathcal{H}}{\delta \rho} = gz$$

• Consider disturbances to a resting basic state $\mathbf{v} = 0$, $\rho = \rho_0(z)$.

$$C = \iiint C(\rho) dx dy dz$$
 with $\frac{\delta C}{\delta \rho} = C'(\rho)$

$$\left. \frac{\delta \mathcal{H}}{\delta \mathbf{u}} \right|_{\mathbf{u} = \mathbf{U}} = -\frac{\delta \mathcal{C}}{\delta \mathbf{u}} \bigg|_{\mathbf{u} = \mathbf{U}} \qquad \frac{\delta \mathcal{H}}{\delta \rho} = gz \qquad \frac{\delta \mathcal{C}}{\delta \rho} = C'(\rho) \qquad C'(\rho_0) = -gz$$

- In the last expression, z must be regarded as a function of ρ_0 via $Z(\rho_0(z)) = z$; so $\rho_0(z)$ must be invertible, i.e. monotonic
- Hence the basic state must be stably stratified

$$C(\rho) = -\int^{\rho} gZ(\tilde{\rho}) d\tilde{\rho}$$

$$A = \iiint \left\{ \frac{\rho_{s}}{2} |\mathbf{v}|^{2} + (\rho - \rho_{0}) gz - \int_{\rho_{0}}^{\rho} gZ(\tilde{\rho}) d\tilde{\rho} \right\} dx dy dz$$

The last two terms in the expression for A can be written

$$-\int_0^{\rho-\rho_0} g[Z(\rho_0+\tilde{\rho})-Z(\rho_0)] d\tilde{\rho} \quad \text{(positive definite)}$$

- Small-amplitude approximation $-\frac{g(\rho-\rho_0)^2}{2(\mathrm{d}\rho_0/\mathrm{d}z)}$ (APE of internal gravity waves)
- Such an APE can be constructed for any Hamiltonian system

Pseudomomentum: In a similar manner, if a basic state u=U
is independent of x (i.e. is invariant with respect to translation
in x), then by Noether's theorem,

$$\partial \mathbf{U}/\partial x = 0$$
 implies $J \frac{\delta \mathcal{M}}{\delta \mathbf{u}} \Big|_{\mathbf{u} = \mathbf{U}} = 0$

which implies $\delta(\mathcal{M} + \mathcal{C}) = 0$ at $\mathbf{u} = \mathbf{U}$ for some Casimir C

 $\mathcal{A} = (\mathcal{M} + \mathcal{C})[\mathbf{u}] - (\mathcal{M} + \mathcal{C})[\mathbf{U}]$ is then both conserved and (pseudomomentum) quadratic in the disturbance

Example: Barotropic flow on the beta-plane

$$\frac{\delta \mathcal{M}}{\delta q} = y, \qquad \frac{\delta \mathcal{C}}{\delta q} = C'(q)$$

• Consider disturbances to an x-invariant basic state $q_0(y)$

$$\delta(\mathcal{M} + \mathcal{C}) = 0$$
 at $q = q_0$ implies $C'(q_0) = -y$

This is analogous to the formula for APE, and similarly,

$$\mathcal{A} = \int \int \left\{ -\int_0^{q-q_0} \left[Y(q_0 + \tilde{q}) - Y(q_0) \right] d\tilde{q} \right\} dx dy$$

where $Y(q_0(y)) = y$, which is negative definite for $dq_0/dy > 0$

- Small-amplitude approximation: $-\frac{(q-q_0)^2}{2(dq_0/dy)}$
- If q_0 is defined to be the zonal mean, then $q_0=\bar{q}, \quad q'=q-\bar{q}$ and the zonal mean of this expression becomes $-\frac{\overline{q'^2}}{2\bar{q}_y}$
- Exactly the same form applies to stratified QG flow, where the negative of this quantity is known as the Eliassen-Palm (E-P) wave activity
- N.B. The sign of this quantity corresponds to the sign of the intrinsic frequency of Rossby waves (negative if $dq_0/dy > 0$)

Relation to wave action: there is a classical result that under slowly varying (WKB), adiabatic conditions, wave action is conserved (Bretherton & Garrett 1969 PRSA)

$$\frac{\partial \hat{A}}{\partial t} + \nabla \cdot \vec{c}_g \hat{A} = 0$$
 where $\hat{A} = \frac{\hat{E}}{\hat{\omega}}$

 $\frac{\partial \hat{A}}{\partial t} + \nabla \cdot \vec{c}_g \hat{A} = 0 \quad \text{where} \quad \hat{A} = \frac{\hat{E}}{\hat{\omega}}$ \hat{E} is the wave energy (always positive definite) and $\hat{\omega}$ is the intrinsic frequency, both measured in the frame of reference moving with the mean flow

- Hence $sgn(\hat{A}) = sgn(\hat{\omega})$
- Under WKB conditions, pseudoenergy and pseudomomentum are related to wave action via $\omega \hat{A}$, $k \hat{A}$ respectively
- However pseudoenergy and pseudomomentum are more general, and extend beyond WKB conditions
 - They require only temporal or zonal symmetry, respectively, in the background state

• Stability theorems: Pseudoenergy and pseudomomentum are conserved in time (for conservative dynamics), and are quadratic in the disturbance (for small disturbances), so for normal-mode disturbances we have

$$\mathcal{A} = \mathcal{A}_0 e^{2\sigma t}$$
 $\sigma \mathcal{A} = 0$ (σ is the real part of the growth rate)

- Then $\mathcal{A} \neq 0$ implies $\sigma = 0$ (normal-mode stability)
- Therefore these conservation laws can provide sufficient conditions for stability/necessary conditions for instability. Indeed, many normal-mode stability theorems (e.g. Pedlosky 1987) result from expressions of the form

$$\sigma \int \{\cdots\} \, \mathrm{d}x = 0$$

where the integral turns out to be just pseudoenergy or pseudomomentum (or some combination of the two)

 Example: Charney-Stern theorem. For stratified QG dynamics, with horizontal boundaries, the pseudomomentum is given by

$$\mathcal{A} = \iiint \rho_0 \left\{ -\int_0^{q-Q} [Y(Q+\tilde{q}) - Y(Q)] d\tilde{q} \right\} dx dy dz$$
$$+ \iiint \left\{ -\int_0^{\lambda_0 - \Lambda_0} [Y_0(\Lambda_0 - \tilde{\lambda}) - Y_0(\Lambda_0)] d\tilde{\lambda} dx dy \right|_{z=0}$$

plus another term with the opposite sign at the top boundary. Here $\Lambda=\Psi_{\tau}$ is proportional to potential temperature.

- Baroclinic instability requires terms of opposite signs so A=0: $Eady \ model$: Interior term vanishes, Λ_y <0 at bottom, Λ_y <0 at top $Charney \ model$: Q_y >0 in interior, Λ_y <0 at bottom $Phillips \ model$: Q_y <0 in lower levels, Q_y >0 in upper levels
- Barotropic instability: can be considered a special case

Other examples of Hamiltonian stability theorems:

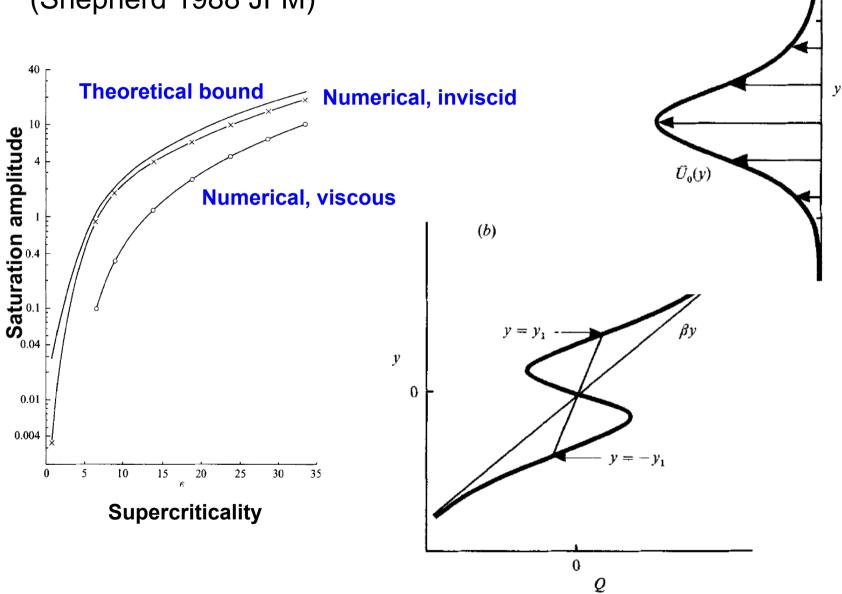
- Static stability, centrifugal stability, symmetric stability
- Rayleigh-Kuo theorem, Fjørtoft-Pedlosky theorem
- Arnol'd's first and second theorems
- Ripa's theorem (shallow-water dynamics)
- Notable exception: stratified shear flow (Miles-Howard theorem)
- These Hamiltonian stability theorems can, in most cases, be generalized to finite-amplitude (Liapunov) stability: i.e. for all ε there exists a δ such that

$$\|\boldsymbol{u}'(0)\| < \delta \qquad \Rightarrow \qquad \|\boldsymbol{u}'(t)\| < \epsilon \quad \forall t$$

 They can also be used to derive rigorous saturation bounds on nonlinear instabilities; e.g. for a statically unstable resting state,

$$\iiint \int \int \frac{1}{2} \rho_0 |\mathbf{v}(t)|^2 dx dy dz \le \mathcal{A}(t) = \mathcal{A}(0) = \iiint APE(0) dx dy dz$$

 Example: Bickley jet on barotropic beta-plane (Shepherd 1988 JFM)



(a)

 Relationship between pseudomomentum and momentum: consider the zonally averaged zonal momentum equation for the barotropic beta-plane:

$$\frac{\partial \bar{u}}{\partial t} = -\frac{\overline{\partial u^2}}{\partial x} - \frac{\overline{\partial u v}}{\partial y} + f\bar{v} - \frac{\overline{\partial p}}{\partial x} = -\frac{\overline{\partial u' v'}}{\partial y}$$

$$\frac{\partial \bar{u}}{\partial t} = -\overline{v'} \left(\frac{\partial u'}{\partial y} - \frac{\partial v'}{\partial x} \right) + \frac{\partial}{\partial x} \left[\frac{1}{2} (u'^2 - v'^2) \right] = \overline{v' q'}$$

The linearized potential-vorticity equation is

$$\begin{split} \frac{\partial q'}{\partial t} + \bar{u} \frac{\partial q'}{\partial x} + v' \frac{\mathrm{d}\bar{q}}{\mathrm{d}y} &= 0 \\ \text{and hence (if } \bar{q}_y \neq 0 \,) \qquad v' &= -\frac{1}{\bar{q}_y} \Big(\frac{\partial q'}{\partial t} + \bar{u} \frac{\partial q'}{\partial x} \Big) \\ \overline{q'v'} &= -\frac{\partial}{\partial t} \Big(\frac{1}{2} \overline{q'^2} \bar{q}_y \, \Big) = \frac{\partial \bar{A}}{\partial t} \quad \text{whence} \quad \frac{\partial \bar{u}}{\partial t} &= \frac{\partial \bar{A}}{\partial t} \end{split} \tag{Taylor identity}$$

 Stratified QG dynamics: zonal-wind tendency equation, temperature tendency equation, and thermal-wind balance together imply

$$\mathcal{L}\left(\frac{\partial \bar{u}}{\partial t}\right) = \frac{\partial^2}{\partial y^2} \overline{\left(v'q'\right)} \quad \text{where} \quad \mathcal{L} = \frac{\partial^2}{\partial y^2} + \frac{1}{\rho_0} \frac{\partial}{\partial z} \frac{\rho_0}{S} \frac{\partial}{\partial z}$$

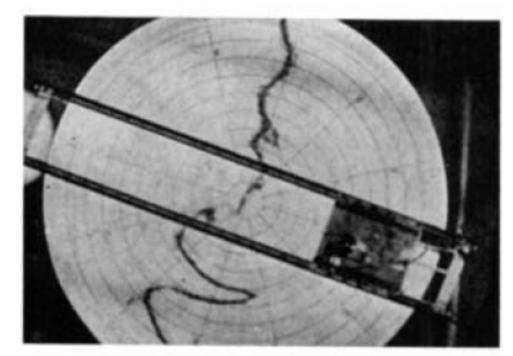
- So it's the same physics, but the zonal-wind response to mixing of potential vorticity is now spatially non-local (the Eliassen balanced response)
- The pseudomomentum conservation law takes the local form (with S being a source/sink)

$$\frac{\partial A}{\partial t} + \nabla \cdot \mathbf{F} = S \qquad \nabla \cdot \mathbf{F} = -\overline{v'q'}$$

$$\mathcal{L}\left(\frac{\partial \bar{u}}{\partial t}\right) = \frac{\partial^2}{\partial y^2} \overline{(v'q')} = -\frac{\partial^2}{\partial y^2} \nabla \cdot \bar{\mathbf{F}} = \frac{\partial^2}{\partial y^2} \left(\frac{\partial \bar{A}}{\partial t} - \bar{S}\right)$$

 So mean-flow changes require wave transience or nonconservative effects (non-acceleration theorem)

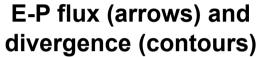
- In the atmosphere, we can generally assume that $q_y > 0$ since q is dominated by β
- Hence A < 0; Rossby waves carry negative pseudomomentum
- Where Rossby waves dissipate, there must be a convergence of negative pseudomomentum, hence a negative torque
- Conservation of momentum implies a positive torque in the wave source region
- This phenomenon is seen in laboratory rotating-tank experiments
- A prograde jet emerges from random stirring, surrounded on either side by retrograde jets (seen in distortion of dye)
 (Whitehead 1975 Tellus)

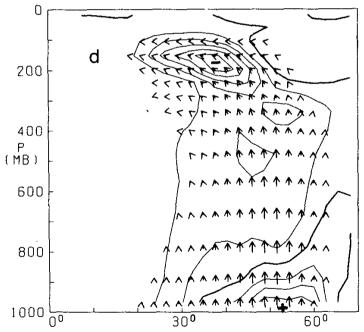


 In the atmosphere, synoptic-scale Rossby waves are generated by baroclinic instability, hence within a jet region

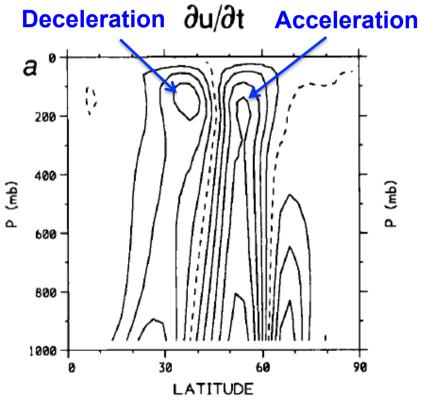
Flux of negative pseudomomentum out of jet corresponds to an upgradient flux of momentum into the jet \(\partial A \) \(\partial \cdots \)

Rossby waves Momentum break & dissipate divergence Momentum Stirring convergence Moméntum Rossby waves divergence break & dissipate Vallis (2006) zonal velocity In fact the wave propagation is up and out (generally equatorward), as seen in these 'baroclinic life cycles' showing baroclinic growth and barotropic decay (Simmons & Hoskins 1978 JAS)





Edmon, Hoskins & McIntyre (1980 JAS)



Haynes & Shepherd (1989 QJRMS)

 The vertical flux of pseudomomentum is expressed in terms of the meridional heat flux

$$-\mathbf{F} = \left(-\rho_0 \overline{u'v'}, (\rho_0 f/\Theta_z) \overline{v'\theta'}\right)$$
 is the Eliassen-Palm (E-P) flux

- Reflects thermal-wind balance: poleward heat flux weakens the thermal wind, accelerating the flow below and decelerating the flow aloft (as in pure baroclinic instability)
- During the wintertime when the stratospheric flow is westerly, stationary planetary Rossby waves can propagate into the stratosphere where they exert a negative torque, acting to weaken the flow from its radiative equilibrium state
- Stationary planetary-wave forcing mechanisms (topography, land-sea temperature contrast) are stronger in the Northern than in the Southern Hemisphere, hence the stratospheric polar vortex is weaker in the Northern Hemisphere

Summary

- Hamiltonian dynamics is applicable to all the important models of geophysical fluid dynamics
 - Provides a unifying framework between various models
 - Systems are infinite-dimensional, and their Eulerian representations are generally non-canonical
 - To exploit Hamiltonian structure all that is needed is to know the conserved quantities of a system
- The most powerful applications are for theories describing disturbances to an inhomogeneous basic state
 - Non-trivial; e.g., wave energy is generally not conserved
 - Useful measures of disturbance magnitude require the use of Casimir invariants, following from Lagrangian invariants
 - Leads to important concepts of pseudoenergy and pseudomomentum: stability theorems immediately follow
 - Important applications are available potential energy and momentum transfer by waves